

# Critical Phenomena

663

1

2<sup>nd</sup> order phase transitions, QFT etc.

3d  $O(N)$  model.

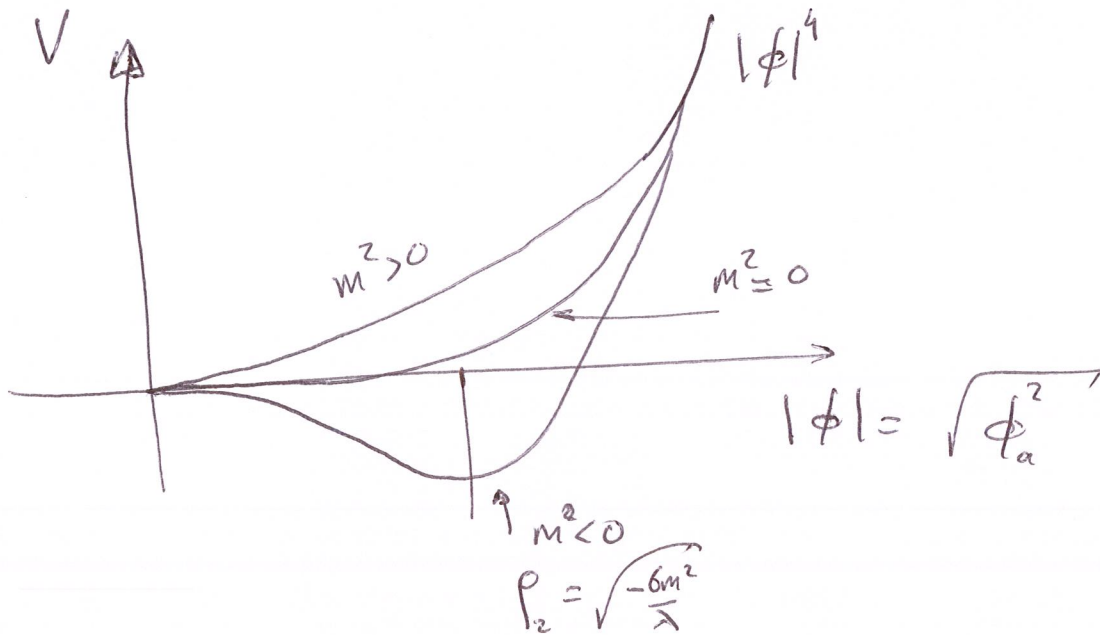
Euclidean

$$\mathcal{L}_E^k = \frac{1}{2} \partial_\mu \phi_a \partial^\mu \phi_a + \frac{1}{2} m^2 \phi_a^2 + \frac{\lambda}{4!} (\phi_a^2)^2$$

$a=1 \dots N$   $\phi_a \in \mathbb{R}$ .

$$V(\phi) = \frac{1}{2} m^2 \phi_a^2 + \frac{\lambda}{4!} (\phi_a^2)^2$$

for  $\partial \phi_a = 0$   
constant field.



$m^2 < 0$

$$\rho = \phi_a^2$$

$$V(\rho) = \frac{1}{2} m^2 \rho + \frac{\lambda}{24} \rho^2$$

$$\frac{\partial V}{\partial \rho} = m^2 + \frac{\lambda}{6} \rho = 0$$

$$\rho_1 = 0 \quad \rho_2 = \sqrt{-\frac{6m^2}{\lambda}}$$

For  $m^2 > 0$   $\langle \phi_a^2 \rangle = 0$  in vacuum.

$\delta\phi_a$  small oscillations around  $\phi_a = 0$

→  $N$  excitations of mass  $m^2$

For  $m^2 = 0$

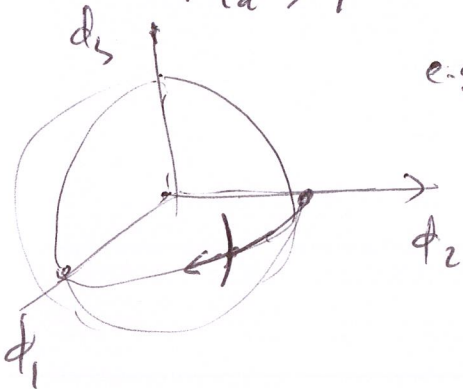
$N$  massless excitations of mass  $m^2$

$$\mathcal{L} = \frac{1}{2} |\partial\phi_a|^2 + \frac{\lambda}{24} (\phi_a^2)^2$$

For  $m^2 < 0$

$\langle \phi_a^2 \rangle \neq 0$   $\langle \phi_a^2 \rangle = -\frac{6m^2}{\lambda}$  in vacuum

e.g.  $N=3$



$N-1$  massless excitations  
(non-linear sigma model on  $S^{N-1}$ )

1 massive excitation ( $\text{rad} > 1$ )

Possible parameterization:

$$\phi_a = R_{ab} \hat{\phi}_b$$

↑ rotation  
↓

$$\hat{\phi}_b = (0, 0, \dots, 0, \hat{\phi})$$

$$\hat{\phi} = \sqrt{-\frac{6m^2}{\lambda}}$$

• However  $SO(N-1)$  does not do anything.

$$SO(N)/SO(N-1)$$

Example

$$N=1 \rightarrow \rho$$

$$N=2 \rightarrow \rho, \vartheta$$

$$N=3 \rightarrow \rho, \underbrace{\vartheta, \varphi}_{\delta^2}$$

$$\begin{aligned} \partial_\mu \phi_a &= \partial_\mu R_{ab} \hat{\phi}_b + R_{ab} \partial_\mu \hat{\phi}_b \\ &= R_{ab} (\partial_\mu \hat{\phi}_b + R_{bc}^{-1} \partial_\mu R_{cd} \hat{\phi}_b) \end{aligned}$$

$$\begin{aligned} R_{ab}^{-1} R_{bc} &= \delta_a^c \\ R_{ab}^{-1} &= R_{ba} \\ R^T &= R^{-1} \end{aligned}$$

$$A_{\mu ab} = R_{ac}^{-1} \partial_\mu R_{cb} = R_{ca} \partial_\mu R_{cb}$$

$$R_{ca} R_{cb} = \delta_{ab}$$

$$\partial_\mu R_{ca} R_{cb} + R_{ca} \partial_\mu R_{cb} = 0$$

$$R_{cb} \partial_\mu R_{ca} = -R_{ca} \partial_\mu R_{cb}$$

$A_\mu$  anti-symmetric.

$$A_{\mu ba} = -A_{\mu ab}$$

$$\vec{\partial}_\mu \vec{\phi} = \partial_\mu \hat{\phi} + A_\mu \cdot \hat{\phi}$$

$$\partial_\mu \phi_a = \partial_\mu \hat{\phi}_a + A_{\mu ab} \hat{\phi}_b$$

$$\begin{aligned} \partial_\mu \phi_a \partial^\mu \phi_a &= \partial_\mu \hat{\phi}_a \partial^\mu \hat{\phi}_a + 2 \partial_\mu \hat{\phi}_a A_{\mu ab} \hat{\phi}_b + \\ &\quad + A_{\mu ab} A_{\mu ac} \hat{\phi}_b \hat{\phi}_c \end{aligned}$$

(4)

$$\vec{\phi}_a = (0, \dots, \rho)$$

$$\rho = \rho_0 + \sigma$$

$$\rho_0 = \sqrt{-\frac{6m^2}{\lambda}}$$

 $A_\mu$  anti-symmetric

$$\partial_\mu \phi_a \partial^\mu \phi_a = \partial_\mu \sigma \partial^\mu \sigma + 2 \cancel{\partial_\mu \sigma A_{\mu NN} \rho_0} + 2 \cancel{\partial_\mu \sigma A_{\mu NN} \sigma} + A_{\mu aN} A_{\mu aN} (\rho_0 + \sigma)^2$$

 $A_{\mu aN}$ : first order in derivatives of the angles.

$$\mathcal{L} = \frac{1}{2} \partial_\mu \sigma \partial^\mu \sigma + \frac{1}{2} A_{\mu aN} A_{\mu aN} \rho_0^2 + \frac{1}{2} m^2 (\rho_0 + \sigma)^2 + \frac{\lambda}{24} (\rho_0 + \sigma)^4 + \frac{1}{2} A_{\mu aN} A_{\mu aN} (2\rho_0 \sigma + \sigma^2)$$

 No potential for the angles  $\rightarrow N-1$  massless fields

$$\frac{1}{2} m^2 \rho_0^2 + m^2 \rho_0 \sigma + \frac{1}{2} m^2 \sigma^2 + \frac{\lambda}{24} (\rho_0^4 + 4\rho_0^3 \sigma + 6\rho_0^2 \sigma^2 + 4\rho_0 \sigma^3 + \sigma^4)$$

$$V_0 = \frac{1}{2} m^2 \rho_0^2 + \frac{\lambda}{24} \rho_0^4$$

vacuum energy.

$$V_1 = m^2 \rho_0 \sigma + \frac{\lambda}{6} \rho_0^3 \sigma = 0 \quad (\rho_0^2 = -6m^2/\lambda)$$

$$V_2 = \frac{1}{2} m^2 \sigma^2 + \frac{\lambda}{4} \rho_0^2 \sigma^2 = \frac{1}{2} m^2 \sigma^2 - \frac{\lambda}{4} \frac{6m^2}{\lambda} \sigma^2 = -\frac{1}{2} m^2 \sigma^2$$

To quadratic order:

$$\mathcal{L} = \frac{1}{2} \partial_n \sigma \partial_n \sigma + \frac{1}{2} \underbrace{(-2m^2)}_{>0} \sigma^2$$

mass  $m_\sigma = \sqrt{-2m^2}$  mass of radial excitation.

$$\mathcal{L} = \frac{1}{2} \rho_0^2 \dot{\mu}_{\text{man}} \dot{\mu}_{\text{man}} \quad (\text{Rememberable only in 2d})$$

$\uparrow$  radius of the sphere  $(f_n)$        $\nwarrow$  kinetic term for motion on the sphere  $S^{N-1}$

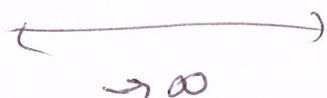
At low energy this dominates  $\rightarrow$  non-linear sigma model.

Cluster decomposition property.

$$\langle 0 | \phi_{a_1}(x_1) \phi_{a_2}(x_2) \phi_{a_{n-1}}(x_{n-1}) \phi_{a_n}(x_n) | 0 \rangle \rightarrow$$

$$\rightarrow \langle 0 | \phi_{a_1}(x_1) - \phi_{a_1}(x_1) | 0 \rangle \langle 0 | \phi_{a_{n-1}}(x_{n-1}) - \phi_{a_n}(x_n) | 0 \rangle$$

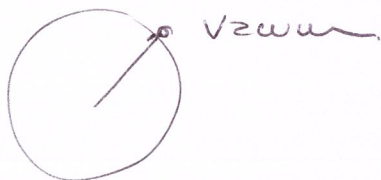
if



If vacuum is a superposition of  $\phi_a$  over sphere then it is not true.

example

10)



$$|0\rangle = \frac{1}{2\pi} \int_{-\pi}^{\pi} d\theta |0\rangle$$

$$\begin{aligned} \langle 0 | \phi_1(x) \phi_1(y) | 0 \rangle &= \frac{1}{2\pi} \int_{-\pi}^{\pi} d\theta \langle 0 | \rho(\theta) \rho(\theta) | 0 \rangle = \\ &= \frac{1}{2\pi} \rho_0^2 \int_{-\pi}^{\pi} d\theta c^2 \theta = \frac{1}{2} \rho_0^2 \end{aligned}$$

$$\langle 0 | \phi_1(x) | 0 \rangle = \frac{1}{2\pi} \int_{-\pi}^{\pi} d\theta \langle 0 | \rho_0(\theta) | 0 \rangle = 0.$$

$$\langle 0 | \phi_1(x) \phi_1(y) | 0 \rangle \neq \langle \phi_1(x) \rangle \langle \phi_1(y) \rangle$$

for any distance  $|x-y| \rightarrow \infty$

for  $|0\rangle$  it is OK

$$\langle 0 | \phi_1(x) \phi_1(y) | 0 \rangle = \rho_0^2 c^2 \theta$$

$$\langle 0 | \phi_1(x) | 0 \rangle = \rho_0 c^2 \theta \quad \checkmark$$

External source

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi_a \partial_\mu \phi_a + \frac{1}{2} m^2 \phi_a^2 + \frac{\lambda}{24} (\phi_a^2)^2 + j_a \phi_a$$

↑  
fixed.

Take  $j_N \neq 0$   $j_{a \neq N} = 0$  (We choose  $N$  axis like this)  
energy minimizes for  $\phi \ll j$

$$\mathcal{L} = \frac{1}{2} \partial_\mu \rho \partial_\mu \rho + \frac{1}{2} A_{\mu\alpha\nu} A_{\mu\alpha\nu} \rho^2 + \frac{1}{2} m^2 \rho^2 + \frac{\lambda}{24} \rho^4 + j_N \rho$$

$m^2 > 0$  phase.  $V(\rho) = \frac{1}{2} m^2 \rho^2 + \frac{\lambda}{24} \rho^4 + j_N \rho$

$$\frac{\partial V}{\partial \rho} = m^2 \rho + \frac{\lambda}{6} \rho^3 + j_N = 0$$

$\rho \approx + j_N / m^2$  small  $j_N$

$m^2 = 0$  phase  $\rho = \left( + \frac{6j_N}{\lambda} \right)^{1/3}$

$m^2 < 0$  phase  $\rho = \rho_0 + \sigma$  small

$$m^2 (\rho_0 + \sigma) + \frac{\lambda}{6} (\rho_0^3 + 3\rho_0^2 \sigma + 3\rho_0 \sigma^2 + \sigma^3) = +j_N$$

$\approx 0$

$$\left[ m^2 + \frac{3\chi}{6} \left( -\frac{6m^2}{\lambda} \right) \right] \sigma = +j_N$$

$$-2m^2 \sigma = +j_N$$

$$\sigma = -\frac{1}{2m^2} j_N$$

For a magnetic system  $\rho \leftrightarrow M$   $j_N \leftrightarrow B$

$$m^2 = f(T) \approx \alpha \frac{(T - T_c)}{T_c} = \alpha t \quad ; \quad t = \frac{T - T_c}{T_c}$$

$$T < T_c \quad m^2 < 0$$

$$T > T_c \quad m^2 > 0$$

Critical exponents.

$$m \sim t^\nu$$

$$\nu = 1/2$$

$$(m^2 \sim t)$$

$$\eta = 0$$

$$\frac{M}{V} \sim |t|^\beta$$

$$M = \langle \rho \rangle = \sqrt{\frac{-6m^2}{\lambda}} \sim |t|^{1/2}$$

$$\beta = 1/2$$

$$\chi \sim |t|^{-\gamma}$$

$$\rho \sim -\frac{j_N}{m^2} = -\frac{j_N}{t}$$

$$\gamma = 1$$

$$\frac{M}{V} \sim H^{1/\delta} \quad t=0$$

$$\rho = \left( -\frac{6j_N}{\lambda} \right)^{1/3}$$

$$\delta = 3$$



$$\langle \sigma(r) \sigma(0) \rangle = \int \frac{d^3 p}{(2\pi)^3} \frac{e^{i\vec{p}\cdot\vec{x}}}{p^2 + m^2} \approx \frac{e^{-mr}}{r}$$

$$\frac{1}{r^p} = \frac{1}{r^{d-2+\eta}}$$

$$\eta = 0$$

$$p = 1$$

$$d = 3$$

$\eta = 0, \nu = 1/2$

$d = 3$

$d = 4$

Relations:

$$\gamma = \nu(2-\eta)$$

$$\beta = \frac{\nu}{2}(d-2+\eta)$$

$$\delta = \frac{d+2-\eta}{d-2+\eta}$$

	1	1
	1/4	1/2
	5	3

$\uparrow$   
wrong.

requires  
fluctuations

$\uparrow$   
OK  
mean field OK

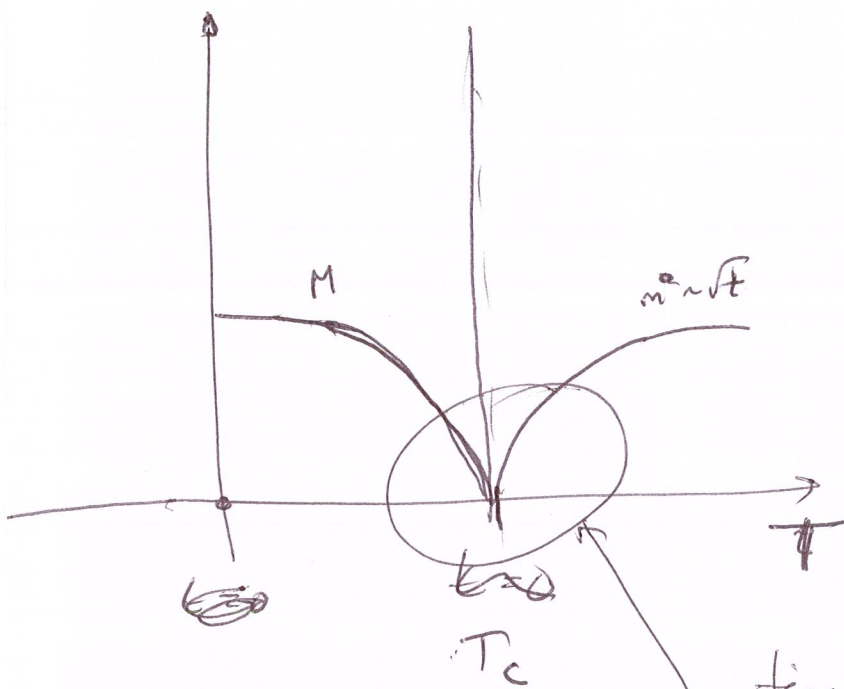
# Fluctuations

Partition function

$$Z = \int \mathcal{D}\phi_a e^{-\int d^3x \mathcal{L}(\phi)}$$

↑ classical partition function or EQFT. Some calculation -  
 Different interpretation. ← Euclidean

Naively <sup>→ to get scale invariance</sup> we should put  $m^2=0$  but there are conditions from fluctuations (thermal or quantum).



critical region dominated by  
 scale invariant theory.

Any QFT has extra dependence on the cut-off.  $\Lambda$ .

At the critical point

$$\langle \phi_a(x) \phi_b(y) \rangle = \delta_{ab} \frac{c \Lambda^{-\eta}}{|x-y|^{d-2+\eta}} = G_{ab}(x,y)$$

$$\int d^d x |\mathcal{D}\phi|^2 \quad M^{-d} \quad M^{+1} \quad M^{-1+d/2}$$

$$[\phi] \sim M^{\frac{d}{2}-1} \quad \uparrow_{\text{mass}}$$

$$\langle d\phi \rangle \sim M^{d-2} \sim L^{2-d} = \frac{1}{L^{d-2}}$$

If we can compute the dependence of  $\langle \phi_a(x) \phi_b(y) \rangle$  on  $\Lambda$  we can get  $\eta$ .

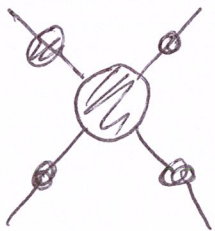
$$\Lambda \frac{\partial}{\partial \Lambda} G_{ab} = \Lambda (-\eta) \frac{1}{\Lambda} G_{ab} = -\eta G_{ab}(x,y)$$

Dependence on cut-off follows from RG.

$$\phi_R = Z_\phi^{-1/2} \phi_B$$

$$\lambda_B = \lambda_B(\lambda_R, \Lambda)$$

If we write everything in terms of renormalized quantities then the cut-off dependence goes away when  $\Lambda \rightarrow \infty$



$$G^{(4)} = \underbrace{G^{(2)} G^{(2)} G^{(2)} G^{(2)}}_n \Gamma^{(4)}$$

$$\Gamma_R^{(n)} = \frac{\langle \phi_R \dots \phi_R \rangle}{\langle \phi \phi \rangle \dots \langle \phi \phi \rangle} = Z_\phi^{n/2} \underbrace{\Gamma_B(g_B(g_R, \Lambda), p_i, \Lambda)}_{\Lambda \rightarrow \infty}$$

is finite (and indep. of  $\Lambda$ )

$$\Lambda \frac{\partial \Gamma_R}{\partial \Lambda} = 0 = \frac{n \Lambda}{2} \frac{\partial Z_\phi}{\partial \Lambda} \frac{1}{Z_\phi^{n/2}} \Gamma_B + Z_\phi^{n/2} \Lambda \frac{\partial \lambda_B}{\partial \Lambda} \frac{\partial \Gamma_B}{\partial \lambda_B} + Z_\phi^{n/2} \Lambda \frac{\partial \Gamma_B}{\partial \Lambda} = 0$$

Ren. conditions  $\Gamma_R^{(4)}(p_i = p_i^R, \mu^R) = \lambda_R \mu^{2\epsilon}$

$\Gamma_R^{(2)}(0) = m_R^2$   $\frac{\partial \Gamma_R^{(2)}(p)}{\partial p^2} = 1$

$$\underbrace{\frac{\eta}{2} \Lambda \frac{\partial \ln Z_f}{\partial \Lambda}}_{-\eta} \Gamma_B + \underbrace{\Lambda \frac{\partial \lambda_B}{\partial \Lambda} \frac{\partial \Gamma_B}{\partial \lambda_B}}_{\beta} + \Lambda \frac{\partial \Gamma_B}{\partial \Lambda} = 0$$

$$\Lambda \frac{\partial \Gamma_B^{(n)}}{\partial \Lambda} + \beta \frac{\partial \Gamma_B^{(n)}}{\partial \lambda_B} - \frac{\eta}{2} \Gamma_B^{(n)} = 0$$

$$\beta \left( \frac{\partial}{\partial \lambda_B} \right) \quad \eta \left( \lambda_B \right) \quad \left( \beta \left( \lambda_B, \frac{\Lambda}{\mu} \right), \text{ but } \mu \text{ does not appear in } \Gamma^{(3)} \right)$$

no units

Suppose  $\beta(\hat{\lambda}_B) = 0$   
 $\leftarrow$  some given value of  $\lambda_B$ .

$$\eta(\hat{\lambda}_B) = \hat{\eta}$$

$$\Lambda \frac{\partial \Gamma_B^{(n)}}{\partial \Lambda} = \frac{\eta}{2} \hat{\eta} \Gamma_B^{(n)} \Rightarrow \Gamma_B^{(n)} = \Lambda^{\eta/2 \hat{\eta}} \underbrace{\tilde{\Gamma}_B^{(n)}}_{\text{indep. of } \Lambda}$$

we get  
cut-off dependence!

$$G_B^{(2)} = (\Gamma^{(2)})^{-1}$$

$$G_B^{(2)} = \Lambda^{-\hat{\eta}} \tilde{G}^{(2)}$$

this is  $\eta$  exponent. critical exp.!!

Example of  $O(n)$

by units so  $\lambda$  no units

$$L_E = \frac{1}{2} \partial_\mu \phi_a \partial_\mu \phi_a + \frac{1}{2} m^2 \phi_a^2 + \frac{\lambda}{4!} (\phi_a^2)^2$$

$$[\phi] = M^{-1+d/2}$$

$$2\epsilon + 4(-1+d/2) = d$$

$$2\epsilon - 4 + 2(4-2\epsilon) = 4-2\epsilon$$

->  $\beta$ -function (  $\beta(d^*) = 0$  )  $\left. \frac{1}{\partial n} \right|_{d^*}$

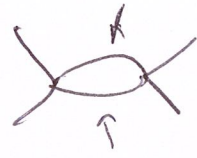
-> critical exponent from combination of operators.

1-loop diagrams.  $(4-2\epsilon \text{ dim})$



$\int \frac{d^4 k}{k^2} \sim \Lambda^2 \rightarrow C_0 p^2 + C_1 + \text{finite piece.}$

actually zero  
divergent  
 $C_0$ : wave-function ren.  
 $C_1$ : mass ren.



$\int \frac{d^4 k}{k^2 k^2} \sim C \ln \Lambda + \text{finite}$

coupling constant ren.

Diagram of a loop with external lines labeled  $a_1 p_1, a_2 p_2, a_3 p_3, a_4 p_4$  and internal momenta  $k, -k+p_1+p_2$ . The loop is labeled  $A^{4\epsilon} \lambda^2$  and  $2 \cdot (4!)^2$ .

	$\phi_{a_1}$	$\phi_{a_2}$	$\phi_{a_3}$	$\phi_{a_4}$	$\int d^4 x \phi_{b_1} \phi_{b_1} \phi_{c_1} \phi_{c_1}$	$\int d^4 x \phi_{b_2} \phi_{b_2} \phi_{c_2} \phi_{c_2}$
$a_2 p_2$	$\square$	$\circ$	$\triangle$	$\times$	$\square$	$\triangle$
$\delta_{a_1 a_2} \delta_{a_3 a_4}$	$\square$	$\circ$	$\triangle$	$\times$	$\square$	$\triangle$
$\delta_{a_1 a_2} \delta_{a_3 a_4}$	$\square$	$\circ$	$\triangle$	$\times$	$\square$	$\triangle$
$\delta_{a_1 a_3} \delta_{a_2 a_4}$	$\square$	$\circ$	$\triangle$	$\times$	$\square$	$\triangle$
$\delta_{a_1 a_4} \delta_{a_2 a_3}$	$\square$	$\circ$	$\triangle$	$\times$	$\square$	$\triangle$

Annotations:  $2 \times 4 \times 4 \times 4 \times 2 \times 2 \times$ ,  $2 \times 4 \times 2 \times 4 \times 2 \times$ ,  $2 \times 4 \times 2 \times 4 \times 2 \times$ ,  $2 \times 4 \times 2 \times 4 \times 2 \times$

$$\delta_{a_1 a_2} \delta_{a_3 a_4} \frac{\lambda^2}{2 \cdot (4!)^2} \cancel{2 \times 4 \times 2 \times 4} (N+4) = \delta_{a_1 a_2} \delta_{a_3 a_4} \lambda^2 \frac{(N+4)}{18}$$

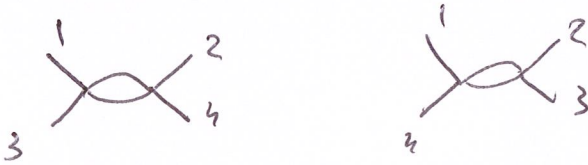
$1 \times 3 \times 1 \times 3$

$$\delta_{a_1 a_3} \delta_{a_2 a_4} \frac{\lambda^2}{2 (1 \times 3)(1 \times 3)} \cancel{2 \times 4 \times 2 \times 4} = \delta_{a_1 a_3} \delta_{a_2 a_4} \frac{\lambda^2}{9}$$

$$\delta_{a_1 a_4} \delta_{a_2 a_3} \frac{\lambda^2}{9}$$

All lines  $B(p_1 + p_2) = \int \frac{d^4 k}{(2\pi)^4} \frac{\Lambda^{2\epsilon}}{k^2 (\epsilon_{p_1 + p_2} - k)^2}$

We add



$$\frac{\lambda^2}{18} B(p_1 + p_2) ((N+4) \delta_{a_1 a_2} \delta_{a_3 a_4} + 2 \delta_{a_1 a_3} \delta_{a_2 a_4} + 2 \delta_{a_1 a_4} \delta_{a_2 a_3})$$

$$\frac{\lambda^2}{18} B(p_1 + p_4) ((N+4) \delta_{a_1 a_4} \delta_{a_2 a_3} + 2 \delta_{a_1 a_3} \delta_{a_2 a_4} + 2 \delta_{a_1 a_2} \delta_{a_3 a_4})$$

$$\frac{\lambda^2}{18} B(p_1 + p_3) ((N+4) \delta_{a_1 a_3} \delta_{a_2 a_4} + 2 \delta_{a_1 a_2} \delta_{a_3 a_4} + 2 \delta_{a_1 a_4} \delta_{a_2 a_3})$$

Bubble in dim. reg.

$$B(p) = \int \frac{d^d k}{(2\pi)^{4-2\epsilon}} \frac{\Lambda^{2\epsilon}}{k^2 (p-k)^2} =$$

$$= \int_0^\infty d\alpha_1 \int_0^\infty d\alpha_2 \Lambda^{2\epsilon} \int \frac{d^d k}{(2\pi)^d} e^{-\alpha_1 k^2} e^{-\alpha_2 (p^2 - 2pk + k^2)}$$

$$= \int_0^\infty d\alpha_1 \int_0^\infty d\alpha_2 \left( \frac{\pi}{\alpha_1 + \alpha_2} \right)^{d/2} \frac{1}{(2\pi)^d} e^{-\frac{\alpha_2 p^2}{\alpha_1 + \alpha_2}} e^{-\alpha_2 p^2} \Lambda^{2\epsilon} e^{-\frac{\alpha_1 \alpha_2}{\alpha_1 + \alpha_2} p^2}$$

$$= \int_0^\infty d\lambda \int_0^\infty d\alpha_1 \int_0^\infty d\alpha_2 \Lambda^{2\epsilon} \frac{\delta(\alpha_1 + \alpha_2 - \lambda)}{(2\pi)^d} \frac{\pi^{d/2}}{(\alpha_1 + \alpha_2)^{d/2}} e^{-\frac{\alpha_1 \alpha_2}{\alpha_1 + \alpha_2} p^2}$$

$$\alpha_1 = \lambda \beta_1, \alpha_2 = \lambda \beta_2$$

$$= \Lambda^{2\epsilon} \int_0^\infty \lambda^2 d\lambda \int_0^\infty d\beta_1 \int_0^\infty d\beta_2 \frac{\delta(\beta_1 + \beta_2 - 1)}{(2\pi)^d \lambda} \frac{\pi^{d/2}}{\lambda^{d/2}} e^{-\lambda \beta_1 \beta_2 p^2}$$

$$= \Lambda^{2\epsilon} \int_0^\infty d\beta_1 \int_0^\infty d\beta_2 \delta(\beta_1 + \beta_2 - 1) \frac{\pi^{d/2}}{(2\pi)^d} \int_0^\infty d\lambda \lambda^{-d/2 + 1} e^{-\lambda \beta_1 \beta_2 p^2}$$

$$= \int_0^1 d\beta_1 \int_0^1 d\beta_2 \delta(\beta_1 + \beta_2 - 1) \frac{\pi^{d/2}}{(2\pi)^d} \Gamma(2 - d/2) (\beta_1 \beta_2 p^2)^{-2 + d/2} \Lambda^{2\epsilon}$$

$$= \frac{\pi^{d/2}}{(2\pi)^d} p^{d-4} \int_0^1 d\beta_1 \beta_1^{d/2-2} (1-\beta_1)^{d/2-2} \Gamma(2-d/2) \Lambda^{2\epsilon}$$



$$\begin{aligned}
 B(p) &= \frac{\pi^{d/2}}{(2\pi)^d} p^{d-4} \Gamma(2-d/2) B(d/2-1, d/2-1) \Lambda^{2\varepsilon} \\
 &= \frac{\pi^{d/2}}{(2\pi)^d} p^{d-4} \Gamma(2-d/2) \frac{\Gamma(d/2-1)\Gamma(d/2+1)}{\Gamma(d-2)} \Lambda^{2\varepsilon} \\
 &= \frac{1}{(4\pi)^{d/2}} p^{d-4} \frac{\Gamma(\frac{4-d}{2}) (\Gamma(d/2-1))^2}{\Gamma(d-2)} \Lambda^{2\varepsilon}
 \end{aligned}$$

$d = 4 - 2\varepsilon$

$$B(p) = \frac{\Lambda^{2\varepsilon}}{(4\pi)^{2-\varepsilon}} p^{-2\varepsilon} \frac{\Gamma(\varepsilon) (\Gamma(1-\varepsilon))^2}{\Gamma(2-2\varepsilon)}$$

$\varepsilon \Gamma(\varepsilon) = \Gamma(1+\varepsilon) \approx 1 + \Gamma'(\varepsilon) \varepsilon$

$\varepsilon \rightarrow 0$

$\Gamma(\varepsilon) \approx 1/\varepsilon$

$B(p) \approx \frac{1}{\varepsilon} \frac{1}{(4\pi)^2}$

pole

$\Lambda$  dependence

$e^{2\varepsilon \ln \Lambda/p}$

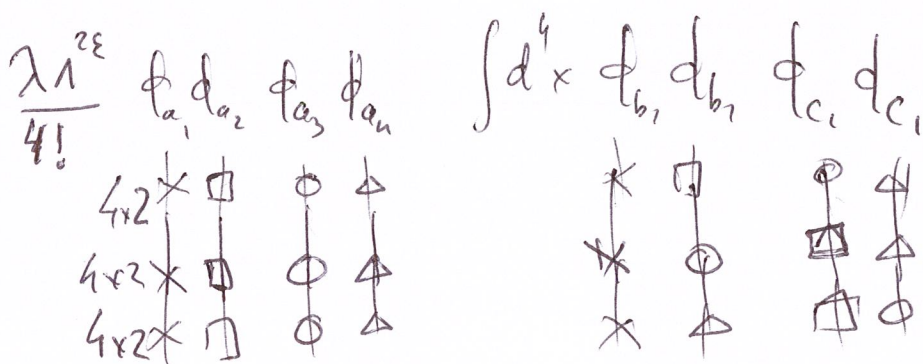
$$B(p) = \frac{1}{\varepsilon} \frac{1}{(4\pi)^2} \left(\frac{\Lambda}{p}\right)^{2\varepsilon} = \frac{1}{16\pi^2 \varepsilon} + \frac{1}{8\pi^2} \ln \Lambda/p + \dots$$

Cut-off dependence.

$$\Gamma_{a_1 a_2 a_3 a_4}^{(4)}(p_1, p_2, p_3, p_4) = \frac{\lambda^2}{18} (N+8) \frac{1}{8m^2} \ln \Lambda (\delta_{a_1 a_2} \delta_{a_3 a_4} + \delta_{a_1 a_3} \delta_{a_2 a_4} + \delta_{a_1 a_4} \delta_{a_2 a_3}) + \text{finite piece}$$

as  $\Lambda \rightarrow \infty$

order  $\lambda$ .



$$\frac{8}{1 \cdot 2 \cdot 3 \cdot 4} \lambda \Lambda^{2\epsilon} (\delta_{a_1 a_2} \delta_{a_3 a_4} + \delta_{a_1 a_3} \delta_{a_2 a_4} + \delta_{a_1 a_4} \delta_{a_2 a_3})$$

$$\Gamma^{(4)} = \frac{\lambda \Lambda^{2\epsilon}}{3} (\delta_{a_1 a_2} \delta_{a_3 a_4} + \delta_{a_1 a_3} \delta_{a_2 a_4} + \delta_{a_1 a_4} \delta_{a_2 a_3})$$

$$\cdot \left[ 1 + \frac{\lambda}{6} (N+8) \frac{1}{8m^2} \ln \Lambda + \dots \right] + \text{finite piece.}$$

$$\mu^{2\epsilon} \lambda_R = \lambda^{2\epsilon} \left( \lambda_B - \frac{\lambda_B^2}{48n^2} (N+8) \ln \frac{d}{\mu} + \text{finite} \right)$$

$$\lambda_R = \cancel{\lambda_B} + \lambda_B \left( \frac{1}{\mu} \right)^{2\epsilon} - \frac{\lambda_B^2}{48n^2} (N+8) \ln \frac{d}{\mu} + \dots$$

$$\stackrel{1\text{-loop}}{=} \lambda_B + \ln \frac{d}{\mu} \left( 2\epsilon \lambda_B - \frac{\lambda_B^2}{48n^2} (N+8) \right) + \dots + \text{finite}$$

$$\lambda_B = \lambda_R - \ln \frac{1}{\mu} \left( 2\epsilon \lambda_R - \frac{\lambda_R^2}{48n^2} (N+8) \right) + \dots$$

$$N \frac{\partial \lambda_B}{\partial N} = -2\epsilon \lambda_R + \frac{\lambda_R^2}{48n^2} (N+8) = \beta$$

$$\beta(\lambda_B) = -2\epsilon \lambda_B + \frac{\lambda_B^2}{48n^2} (N+8)$$

$\lambda_B = 0$  scale inv.

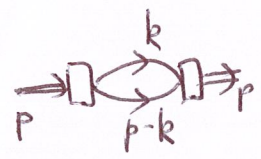
$$\lambda_B^* = \frac{96\epsilon n^2}{(N+8)}$$

$$g = \frac{\lambda}{16n^2}$$

$$g_B^* = \frac{6\epsilon}{N+8}$$

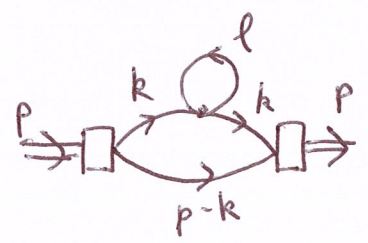
zero of  $\beta$ -function

$$\langle \phi_a^2 \phi_b^2 \rangle$$

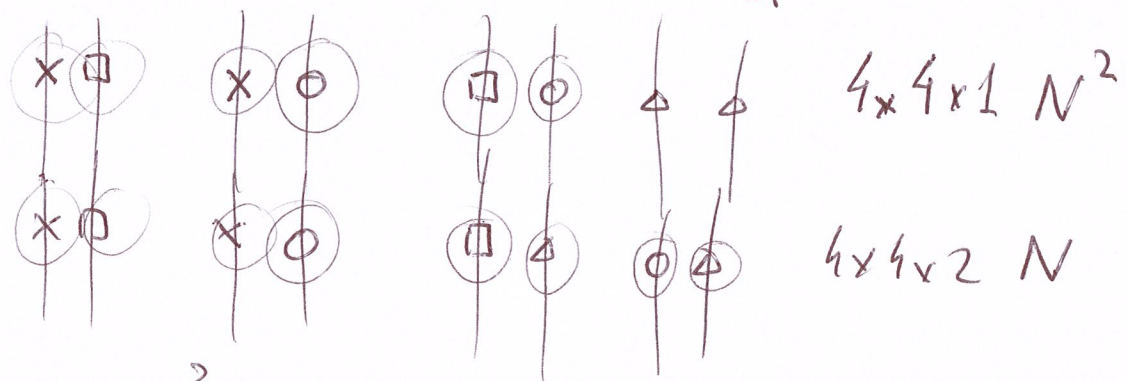


$$\phi_a \phi_a \quad \phi_b \phi_b$$

$$\times \quad \phi \quad \times \quad \phi \quad 2N B(p)$$



$$-\frac{\lambda}{4!} \phi_a \phi_a \quad \phi_b \phi_b \quad \phi_{a_1} \phi_{a_1} \quad \phi_{b_1} \phi_{b_1}$$

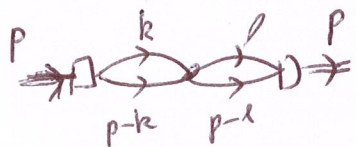


$$-\frac{\lambda}{1234} \times \times N(N+2) \int \frac{d^d k}{(2\pi)^d} \int \frac{d^d l}{(2\pi)^d} \frac{1}{(k^2)^2} \frac{1}{(p-k)^2} \frac{1}{l^2}$$

$$\int \frac{d^d l}{l^2} \rightarrow \infty \quad \text{dim. reg.}$$

$$-\frac{2\lambda N}{3} (N+2) \Delta(0) \int \frac{d^d k}{(2\pi)^d} \frac{1}{(k^2)^2} \frac{1}{(p-k)^2}$$

$$\text{A div} = -3$$

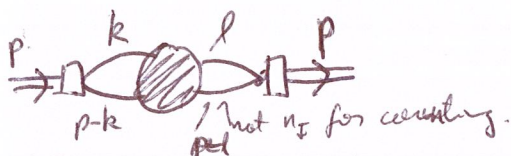


$$-\frac{\lambda}{4!} \phi_a \phi_a \phi_b \phi_b \phi_{a_1} \phi_{a_1} \phi_{b_1} \phi_{b_1}$$

$$-\frac{\lambda}{(1 \times 2 \times 3 \times 4)} \delta (N+2) N B^2(p)$$

$$\langle \phi_a^2 \phi_b^2 \rangle = 2N B(p) - \frac{\lambda N (N+2)}{3} B^2(p) + \dots$$

$$= 2N B(p) \left( 1 - \frac{\lambda}{6} (N+2) B(p) + \dots \right)$$



$$4n_v = 2n_I + 4$$

$$n_I = 2n_v - 2$$

$$n_L = n_I + 2 - n_v + 1$$

$$n_L = 2n_v - 2 - n_v + 1 = n_v + 1$$

$$\# \text{div} = d n_L - 2(n_I + 4) = d n_v + d - 4n_v + 4 - 8 = (d-4)n_v + d - 4$$

$$\# \text{div} = (n_v + 1)(d-4)$$

$$d=4 \quad \# \text{div} = 0 \quad \Rightarrow \quad \underline{\ln} \Lambda \times \text{Constant}$$

$\Gamma^{(2,0)}$  has a divergence even for the free theory.  
we need to subtract

$$\Gamma_R^{(2,0)}(p) = Z_2^{-2} Z_f^{-2} \left( \Gamma_B^{(2)}(p) - \underbrace{\Gamma_B^{(2)}(\mu)}_{\text{cancel this.}} \right)$$

Take also  $\Gamma_R^{(2,0)}(p=3\mu) = -1$  for example.

$$\Gamma_R^{(2,0)}(p=\mu) = 0$$

$$\Gamma_B^{(2)}(p) - \Gamma_B^{(2)}(\mu) = 2N(B(p) - B(\mu)) - \frac{\lambda N(N+2)}{3} (B^2(p) - B^2(\mu)) =$$

$$= 2N(B(p) - B(\mu)) \left[ 1 - \frac{\lambda(N+2)}{6} (B(p) + B(\mu)) \right]$$

$$B(p) = \frac{1}{8\pi^2} \ln \frac{1}{p} + b \quad \uparrow \text{finite constant.}$$

$$B(p) - B(\mu) = \frac{1}{8\pi^2} \ln \frac{\mu}{p} = \text{finite } \checkmark.$$

$$B(p) + B(\mu) = \frac{1}{8n^2} \left( \ln \frac{1}{p} + \ln \frac{1}{\mu} \right) + 2b \quad (4)$$

$$= \frac{1}{4n^2} \ln \frac{1}{\sqrt{\mu p}} + 2b$$

$$\Gamma_R^{(2,0)}(p) = Z_2^2 Z_f^{-2} \frac{2N}{8n^2} \ln \frac{1}{p} \left( 1 - \frac{\lambda(N+2)}{24\pi^2} \ln \frac{1}{\sqrt{\mu p}} + \tilde{b} \right)$$


$$\Gamma_R^{(2,0)}(2\mu) = Z_2^2 Z_f^{-2} \frac{N}{4n^2} \ln \left( \frac{1}{2} \right) \left( 1 - \frac{\lambda(N+2)}{24n^2} \ln \frac{1}{\sqrt{2}\mu} + \tilde{b} \right) = -1$$

$$Z_f = 1 + \mathcal{O}(\lambda^2)$$

$$Z_2^2 = \sqrt{\frac{4n^2}{N \ln \frac{1}{2}}} \left( 1 + \frac{\lambda(N+2)}{48n^2} \ln \frac{1}{\sqrt{2}\mu} \right)$$

$$\wedge \frac{\partial \ln Z_2}{\partial n} = \frac{\lambda(N+2)}{48n^2} = \frac{g^*(N+2)}{3} = \eta_2$$

$$Z_2 \sim \left( \frac{\Lambda}{\mu} \right)^{\frac{1}{3} g^*(N+2)}$$



$$\langle \phi_B^\dagger \phi_B^\dagger \rangle = \frac{C \Lambda^{-2\Delta_\phi^2 + 2d - 4}}{|x-y|^{2\Delta_\phi^2}}$$

$$[f] = M^{d/2 - 1}$$

$$\int e^{ip(x-y)} d^d x \frac{1}{|x-y|^{2\Delta_\phi^2}} = P^{2\Delta_\phi^2 - d}$$

$$\int_B \langle \phi_B^\dagger \rangle = \frac{\tilde{c} \Lambda^{-2\Delta_\phi^2 + 2d - 4}}{p^{d - 2\Delta_\phi^2}} = \frac{\tilde{c} \left( \Lambda^{-\Delta_\phi^2 + d - 2} \right)^2}{p^{d - 2\Delta_\phi^2}}$$

$$= Z_2^{-2} \sqrt{R}^{(2)}(p^2)$$

$$Z_2^{-2} = \left( \frac{\Lambda}{\mu} \right)^{-\frac{2}{3} g^* (N+2)}$$

$$-\Delta_\phi^2 + d - 2 = -\frac{1}{3} g^* (N+2)$$

$$\Delta_\phi^2 = d - 2 + \frac{1}{3} g^* (N+2)$$



$$S = S_{\text{CFT}} + \alpha \int d^d x \, t \, \phi^2$$

$$= S_{\text{CFT}} + \tilde{\alpha} \int d^d x \, \xi^{-1/\nu} \phi^2$$

$\downarrow$   
 $-\Delta_\phi^2$   
 $\xi$

$$d - \frac{1}{\nu} - \Delta_\phi^2 = 0$$

$$\frac{1}{\nu} = d - \Delta_\phi^2 = d - d + 2 - \frac{g^*}{3} (N+2)$$

$$\nu = \frac{1}{2 \left( 1 - \frac{g^*}{3} (N+2) \right)} \approx \frac{1}{2} \left( 1 + \frac{g^*}{3} (N+2) \right)$$

$$\nu = \frac{1}{2} \left( 1 + \frac{g^* \epsilon}{N+8} \frac{N+2}{\epsilon} \right)$$

$$= \frac{1}{2} \left( 1 + \frac{\epsilon}{N+8} \frac{N+2}{\epsilon} \right)$$