EAPS 53600: Introduction to General Circulation of the Atmosphere Spring 2020 NASA GEOS-5 Computer Model

#### Topic: Madden-Julian Oscillation (MJO)

Reading: 1. Hartmann Ch 8.2.2

Prof. Dan Chavas

White: total precipitable water (brigher white = more water vapor in column) Colors: precipitation rate ( $0 - 15 \frac{mm}{hr}$ , red=highest)



Source: https://svs.gsfc.nasa.gov/cgi-bin/details.cgi?aid=30017

Learning outcomes for today:

- Describe what the Madden-Julian Oscillation is and how to identify it
- Explain why it exists
- Explain what causes it to propagate eastward



## Madden Julian Oscillation (MJO) Basics





https://www.youtube.com/watch?v=UsWHHE\_jkGE



https://www.youtube.com/watch?v=V3HE7WMm-KM



Duration:  $\tau \sim 1 - 2$  months Frequency: 1 - 2 months

- A pulsing of organized enhanced rainfall that begins over the Indian Ocean
- Driven by internal atmospheric feedbacks – only starting to be better understood!



of-US-CLIVAR fig1 255244827



Hartmann Fig 8.7

 $\Delta t \approx 60 \ days$ 

 $u_{t.MIO} \approx$ 



#### Negative OLR anomaly (enhanced deep convective clouds) Positive OLR anomaly (suppressed deep convective clouds)

Contour interval =  $2 \text{ W/m}^2$ 

#### Wind vector and OLR associated with 1<sup>st</sup>/2<sup>nd</sup> EOFs of 200 hPa and 850 hPa zonal wind, Oct-Mar



#### Second EOF (~Phase 4-5)



Wheeler and Hendon (2004) Hartmann Fig 8.8

More about EOFs and their prosterns: https://journals.ametsoc.org/doi/pdf/10.1175/2009JCLI3062.1

#### Real-time daily MJO tracking Wheeler and Hendon (2004) MJO phase diagrams

Phase and amplitude of the MJO based on the principal component time series of the leading two EOFs from a combined EOF analysis using 850 hPa zonal wind, 200 hPa zonal wind and OLR.

Counter-clockwise movement around the diagram indicates an eastward propagating signal across eight phases from the Indian Ocean to the Pacific and later the western hemisphere.

Color of lines distinguish different months and dates are annotated.

The farther away from the center of the circle the stronger the MJO signal.

https://www.cpc.ncep.noaa.gov/product s/precip/CWlink/MJO/whindex.shtml





# Physics



#### If given enough time (~10s of days), convection spontaneously self-aggregates in uniform non-rotating radiative-convective equilibrium

### Small domain: $L_{x,y}$ = 198 km (Regular convection)

t=0 minutes Clouds (white surfaces), surface temperature (colors)

Small domain (L=198km): disorganized convection



### Large domain: $L_{x,y} = 510$ km (Aggregated convection)

Large domain (L=510km): convection self-aggregates



#### What creates the MJO and why does it move eastward?

#### **Khairoutdinov and Emanuel (2018, JAS)** How an MJO forms and moves an aquaplanet without SST gradients

Recent studies have suggested that the Madden–Julian oscillation is a result of an instability driven mainly by cloud–radiation feedbacks, similar in character to self-aggregation of convection in nonrotating, cloud-permitting simulations of radiative–convective equilibrium (RCE). Here we bolster that inference by simulating radiative–convective equilibrium states on a rotating sphere with constant sea surface temperature, using the cloud-permitting System for Atmospheric Modeling (SAM) with 20-km grid spacing and extending to walls at 46° latitude in each hemisphere. Mechanism-denial experiments reveal that cloud–radiation interaction is the quintessential driving mechanism of the simulated MJO-like disturbances, but wind-induced surface heat exchange (WISHE) feedbacks are the primary driver of its eastward propagation. WISHE may also explain the faster Kelvin-like modes in the simulations. These conclusions are supported by a linear stability analysis of RCE states on an equatorial beta plane.

Khairoutdinov, M.F. and Emanuel, K., 2018. Intraseasonal variability in a cloud-permitting near-global equatorial aquaplanet model. *Journal of the Atmospheric Sciences*, 75(12), pp.4337-4355.





FIG. 2. Snapshots of column-integrated water vapor on (top) day 15 and (bottom) day 280.



#### Moist region (enhanced convection)



**Precipitable Water Anomalies** 

This is <u>one</u> continuous propagating disturbance!

 $au_{circumnavigate} pprox 60 \ days$  $u_t \ pprox +8 \ m/s$ 

FIG. 4. Hovmöller diagram of equatorial (10°S–10°N) precipitable water anomalies in the CTRL simulation for (a) unfiltered and (b) only zonal wavenumber-1 mode remaining.





FIG. 5. Wavenumber–frequency (left) symmetric and (right) antisymmetric spectra about the equator of OLR over the equatorial belt (10°S–10°N). The ratio of the total power over the background power is shown. The thin lines represent the 12-, 25-, and 50-m equivalent depth of shallow-water theory.

#### MJO spectral signature



FIG. 7. Anomalies for the composite MJO-like disturbance with zonal wavenumbers larger than 4 filtered out for (a) precipitation rate, (b) column MSE, and relative vorticity with corresponding wind at (c) 850 and (d) 200 hPa.







#### Composite MJO: Column Moist Static Energy Tendencies



FIG. 8. Composite budget of the column-integrated MSE tendencies for the composite MJO-like disturbance in the CTRL simulation due to (a) surface enthalpy flux, (b) net radiation heating, (c) advection of MSE, and (d) net local tendency due to all processes. The center of MJO-like disturbance is marked by a dot.

#### Composite MJO: Column Moist Static Energy Tendencies





#### Budget equation for column-integrated MSE

$$\frac{\partial \langle m \rangle}{\partial t} = -c \frac{\partial \langle m \rangle}{\partial x} - \langle \mathbf{V} \cdot \nabla m \rangle - \left\langle \omega \frac{\partial m}{\partial p} \right\rangle + F_0 + R_0 - R_t.$$
(2)
Net radiative
Advection
Surface
enthalpy
fluxes

1) Net radiative heating increases  $\langle m \rangle$  over MJO center (where  $\langle m \rangle$  is already large) – a positive feedback that maintains the MJO. Physically: more clouds/moisture = stronger greenhouse effect

2) Surface fluxes increase  $\langle m \rangle$  to the east of the MJO – a feedback that causes it to propagate eastward. Physically: surface fluxes increase with wind speed, and surface wind speeds are strongest to the strong (easterly inflow + background equatorial easterlies)

#### "Mechanism-denial experiments"



#### "Mechanism-denial experiments"



FIG. 9. Hovmöller diagrams of equatorial (10°S–10°N) zonal wavenumber-1-filtered OLR anomalies in CTRL and various mechanismdenial experiments (see text for acronyms).

What if you do not allow radiative feedbacks? (By setting  $\dot{Q}(z) = \dot{Q}(z)$  constant everywhere, Weaker MJO! Similar translation speed



FIG. 11. Hovmöller diagrams of equatorial (10°S–10°N) zonal wavenumber-1-filtered OLR anomalies in the CTRL case interrupted at day 280 by homogenization of the surface wind speed in surface enthalpy flux computations (similar to the NO\_WISHE case).

#### More proof that WISHE is really fundamental.

Instantaneously turn off wind-speed dependence of surface fluxes. MJO stops moving, nearly disappears.



The MJO is not itself a unique wave mode.

It can look like a wave because it propagates. But it's probably better thought of simply as a coherent disturbance.

It is a disturbance that emerges from slow (O(10 day)) feedbacks between moisture, radiation, and circulation.



### For more information + a remarkable *analytical* solution for this MJO-like disturbance

Khairoutdinov, M.F. and Emanuel, K., 2018. Intraseasonal variability in a cloud-permitting near-global equatorial aquaplanet model. *Journal of the Atmospheric Sciences*, 75(12), pp.4337-4355.



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