EAPS 53600: Introduction to General Circulation of the Atmosphere Spring 2020 Prof. Dan Chavas

Topic: The Transformed Eulerian

VallisE Ch 9.3-9.4, Ch 12.1

Mean (TEM)

Reading:

NASA GEOS-5 Computer Model

White: total precipitable water (brigher white = more water vapor in column) Colors: precipitation rate $(0 - 15 \frac{mm}{hr})$, red=highest)



Source: https://svs.gsfc.nasa.gov/cgi-bin/details.cgi?aid=30017

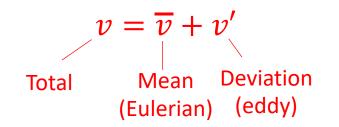
Learning outcomes for today:

- Describe how the Transformed Eulerian Mean differs from a regular Eulerian mean
- Explain the relationship among wave activity, EP fluxes, EP flux divergence/convergence, and acceleration/deceleration of the zonal-mean flow
- Describe how these concepts are relevant to our real atmosphere



Eulerian mean: the normal way you take an average, think about eddies

(average at a fixed location in time (e.g. annual mean) and/or space (e.g. zonal mean))



Transformed Eulerian mean (TEM): a modified framework that lets you directly link eddy effects to changes in mean zonal flow

 Moves eddy PV flux effects out of thermodynamic equation and entirely into momentum equation



Last time: The eddy meridional PV flux can be written as the EP flux divergence

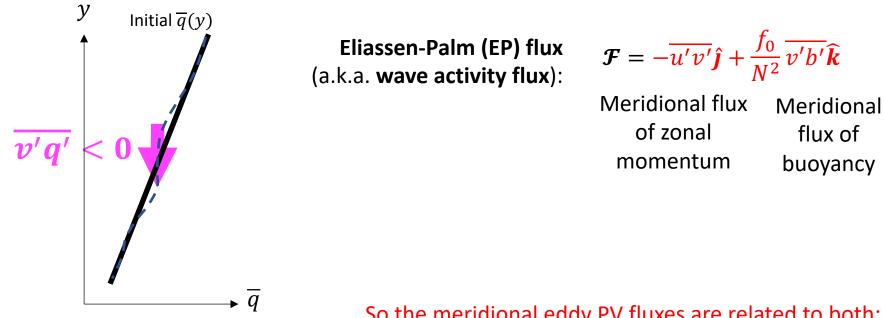
Eddy meridional PV flux:

$$\overline{\nu' q'} = \nabla \cdot \mathcal{F} \tag{9.28}$$

flux of

buoyancy

(9.27)



So the meridional eddy PV fluxes are related to both:

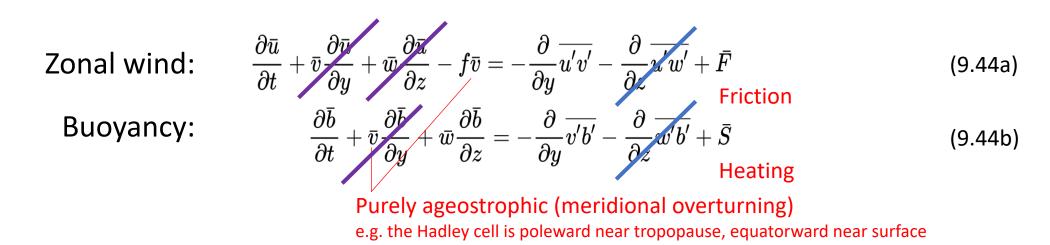
- meridional fluxes in zonal wind (u)1)
- 2) meridional fluxes in buoyancy (b)

Which of these do eddies actually do? Or is it both? Let's go back to the individual equations for each...



Eddy PV flux is actually an eddy momentum flux + eddy buoyancy flux

Eulerian mean zonal-mean equations, Boussinesq fluid:



Quasi-geostrophic:

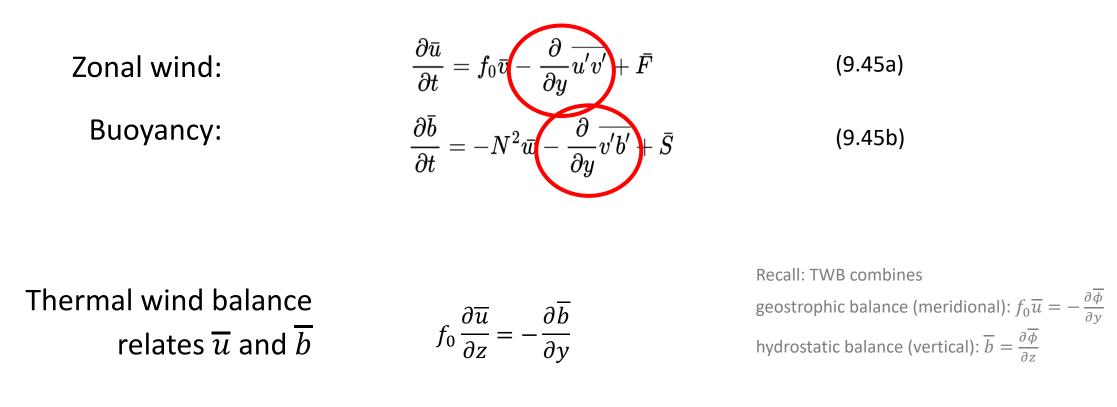
neglect vertical eddy flux convergences (terms with w')

neglect ageostrophic velocities (\overline{v} or \overline{w}) except when multiplied by f_0 or N^2

Note:
$$\frac{\partial \overline{b}}{\partial z} = N^2$$



Eddy PV flux is actually an eddy momentum flux + eddy buoyancy flux

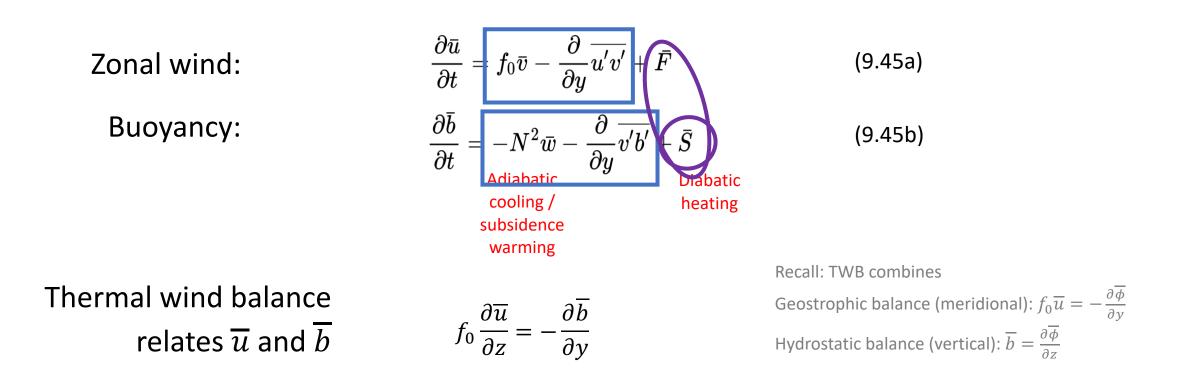


Now we can see the individual eddy fluxes. However, because of thermal wind balance...

- eddy momentum fluxes can change \overline{u} , which via TWB changes \overline{b}
- eddy buoyancy fluxes can change b, which via TWB changes \overline{u}

So we still don't actually know how the eddy fluxes change \overline{u} and b!

Eddy PV flux is actually an eddy momentum flux + eddy buoyancy flux



In the extratropics: these are the dominant balance and these are typically small residuals

S: small but <u>important</u> -- this drives the mean meridional overturning (note: in the tropics it creates the temperature gradients that drive the Hadley cell)

> Can we **transform these equations** into a more useful form, that **accounts for this residual term** \overline{S} ?



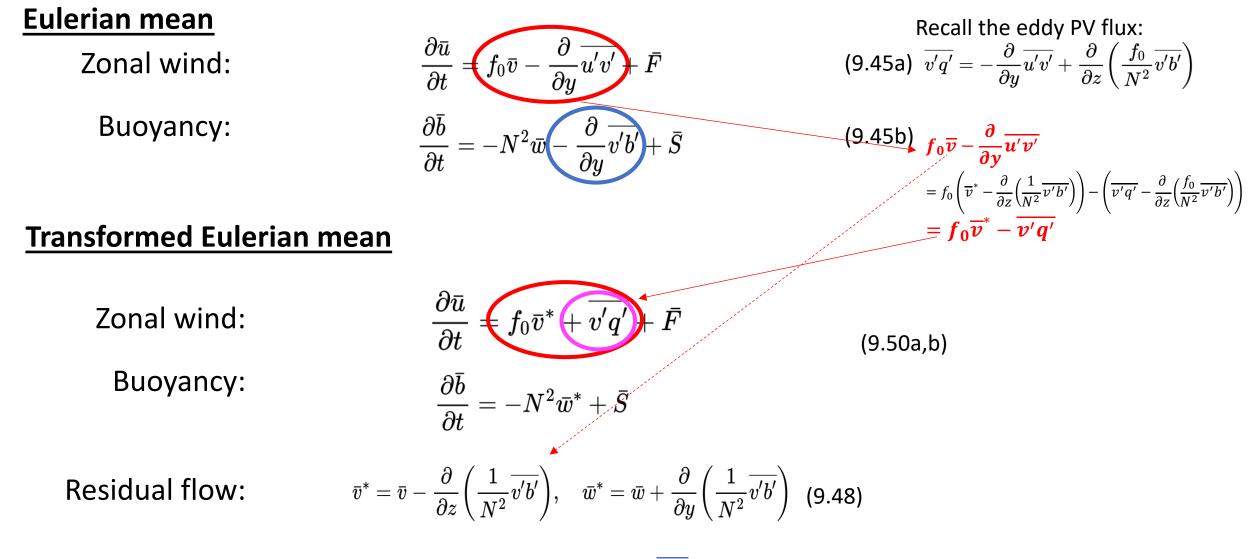
Let's absorb the dominant thermodynamic balance into our equations

 $rac{\partial ar{u}}{\partial t} = f_0 ar{v} - rac{\partial}{\partial u} \overline{u'v'} + ar{F}$ Zonal wind: (9.45a) Dominant balance: $\overline{w}_{bal} = \frac{\partial}{\partial v} \left(\frac{1}{N^2} \overline{v'b'} \right)$ **Buoyancy:** $rac{\partial ar{b}}{\partial t} = -N^2 ar{w} - rac{\partial}{\partial u} \overline{v'b'} + ar{S}$ (9.45b) Define residual flow: $\overline{w}^* = \overline{w} - \overline{w}_{bal}$ Mass continuity: $egin{aligned} egin{aligned} ar v,ar w) &= \left(-rac{\partial\psi_m}{\partial z},rac{\partial\psi_m}{\partial y}
ight) \end{aligned}$ ψ_m : Eulerian mean (9.46)relates \overline{v} and \overline{w} meridional streamfunction where $\frac{\partial \overline{v}}{\partial y} + \frac{\partial \overline{w}}{\partial z} = 0$ Plug in \overline{w}_{bal} to get \overline{v}_{bal} $\frac{\partial \overline{v}_{bal}}{\partial y} = -\frac{\partial}{\partial z} \frac{\partial}{\partial y} \left(\frac{1}{N^2} \overline{v'b'}\right) = \frac{\partial}{\partial y} \left(-\frac{\partial}{\partial z} \left(\frac{1}{N^2} \overline{v'b'}\right)\right)$ $\bar{v}^* = \bar{v} - \frac{\partial}{\partial z} \left(\frac{1}{N^2} \overline{v'b'} \right), \quad \bar{w}^* = \bar{w} + \frac{\partial}{\partial y} \left(\frac{1}{N^2} \overline{v'b'} \right) \quad (9.48) \qquad \overline{v}_{bal} = -\frac{\partial}{\partial z} \left(\frac{1}{N^2} \overline{v'b'} \right)$ Define residual flow: $\overline{v}^* = \overline{v} - \overline{v}_{hal}$ where $\frac{\partial \overline{v}^*}{\partial y} + \frac{\partial \overline{w}^*}{\partial z} = 0$ (Mass continuity has to apply to the residual, too!) with **residual** $(ar{v}^*,ar{w}^*)=\left(-rac{\partial\psi^*}{\partial z},rac{\partial\psi^*}{\partial y}
ight)$ (9.47b) ψ^* : <u>Transformed</u> Eulerian each (9.47a) (residual) meridional streamfunction $\psi^* \equiv \psi_m + rac{1}{N^2} \overline{v'b'}$

streamfunction ψ^* :

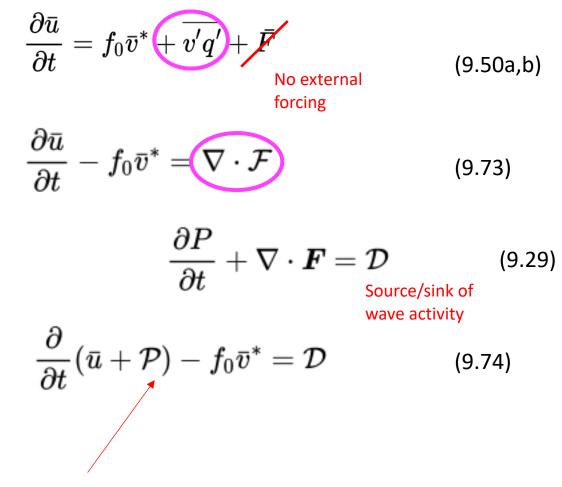
Define a **residual mean** meridional circulation ($\overline{v}^*, \overline{w}^*$):

Let's absorb the dominant thermodynamic balance into our equations



We've effectively moved the eddy buoyancy flux divergence term $(\overline{v'b'})$ into the momentum equation and combined it with the zonal momentum flux divergence term $(\overline{u'v'})$. The result? A <u>single</u> eddy forcing term: the eddy PV flux, $\overline{v'q'}$, that <u>directly</u> changes the zonal wind $(\frac{\partial \overline{u}}{\partial t})!$

Now let's link back to wave activity and the EP flux

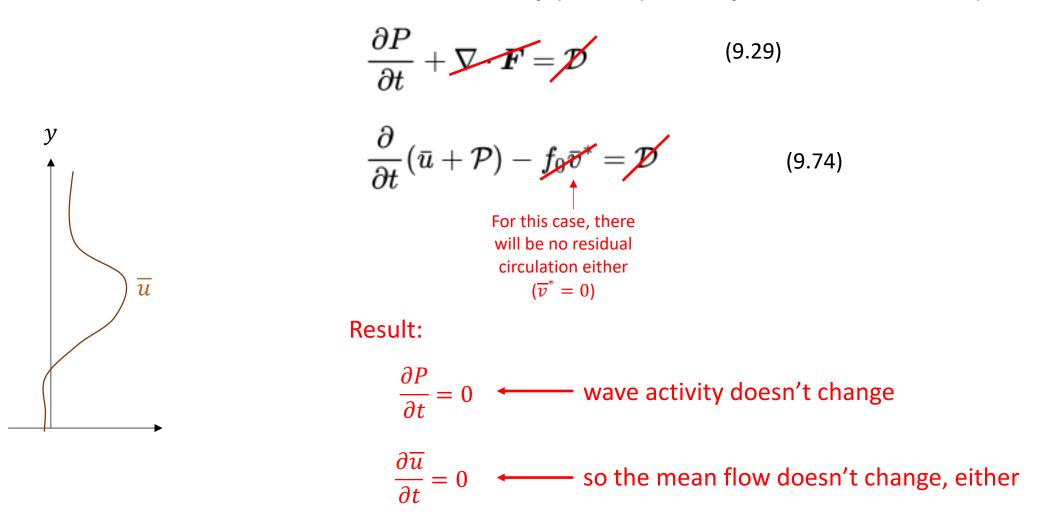


Recall that the wave activity, \mathcal{P} , is also called "**pseudomomentum**". This is why: \mathcal{P} acts dynamically like a zonal-mean zonal momentum per unit mass (and it has units of m/s)

We can use these equations to understand a lot about how the mean flow is affected by eddies. Let's think about some examples...



Case 1: No sources/sinks of wave activity ($\mathcal{D} = 0$), steady motionless waves ($\nabla \cdot \mathcal{F} = 0$)

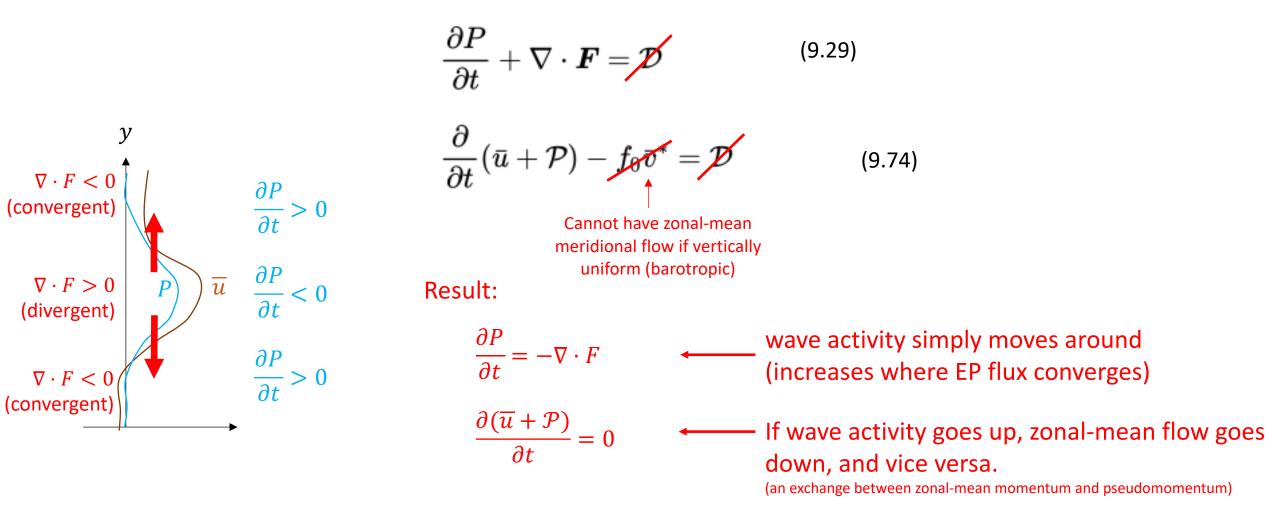


This is called the "non-acceleration" result. The mean flow is not accelerated at all, regardless of where wave activity is found.



Ok, so waves alone aren't enough to change the mean flow. What if waves could move around?

Case 2: No sources/sinks of wave activity ($\mathcal{D} = 0$), barotropic fluid ($\overline{\nu} = 0$)

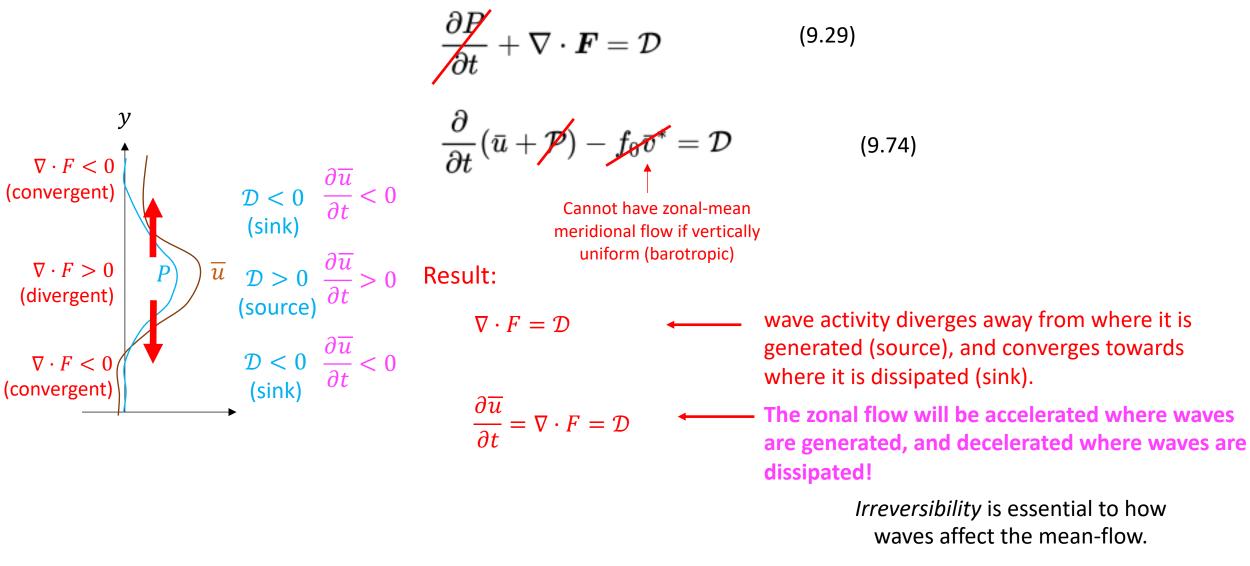


But this is <u>not realistic</u>: something has to generate (and dissipate) this wave activity in the first place...

Why does this matter?

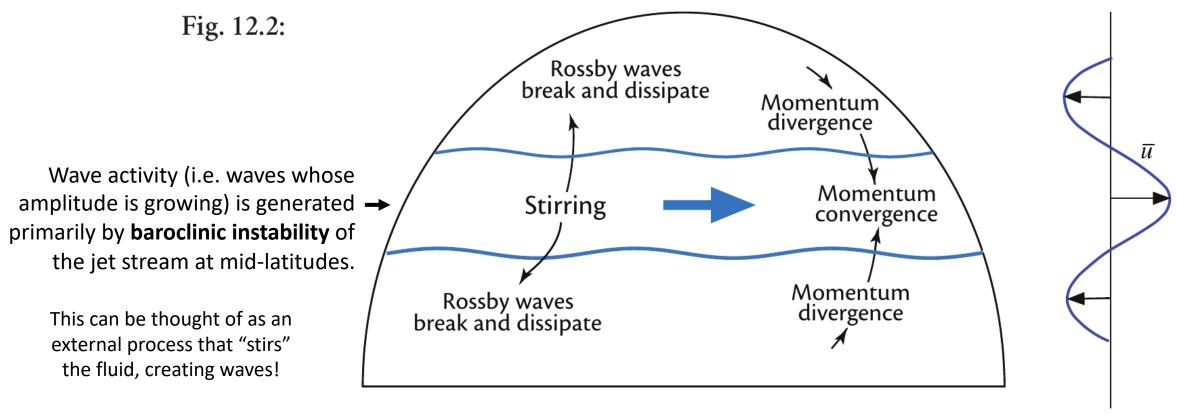
It means that **sources/sinks of wave activity** are **fundamental to changing the (barotropic) mean flow.**

Case 3: steady wave activity, barotropic fluid ($\overline{v} = 0$)



Ok, so how does this manifest itself in our actual atmosphere?





Zonal flow

In this way, then waves are converging zonal momentum toward their source region. This momentum is taken from lower/higher latitudes where the waves break and dissipate.

In essence, then, baroclinic instability of the mid-latitude (barotropic) jet helps to sustain the jet's very existence. This instability tries to break down the zonal-mean jet into eddies *and* restore its zonal-mean state *simultaneous* Can we apply these concepts to understand the large-scale circulation of our atmosphere?

Where do real eddies transport momentum vs. heat (buoyancy)?

How does the global momentum budget fit into this?

Where does the Ferrel cell come from?

We'll cover these topics in the next lecture.

A couple of notes:

- A nice summary of the TEM equations is provided in the box VallisE p. 186
- Why would a meridional buoyancy flux (v'b') act like a force (^{∂u}/_{∂t})? This seems weird/magical. (Recall: this shows up as the vertical component of the EP flux) See VallisE Section 9.4.3 the buoyancy flux acts like a form drag created by sloping interfaces that places a stress on the fluid (i.e. a sink of momentum). Transporting heat (buoyancy) acts to change the slope of pressure surfaces and thus changes this form drag.



Now go to Blackboard to answer a few questions about this topic!

